



## Corrigendum

## Corrigendum to “Stable chromium isotopic composition of meteorites and metal–silicate experiments: Implications for fractionation during core formation” [Earth Planet. Sci. Lett. 435 (2016) 14–21]



P. Bonnand<sup>a,\*</sup>, H.M. Williams<sup>b</sup>, I.J. Parkinson<sup>c,d</sup>, B.J. Wood<sup>a</sup>, A.N. Halliday<sup>a</sup>

<sup>a</sup> Department of Earth Sciences, University of Oxford, South Parks Road, Oxford, OX1 3AN, United Kingdom

<sup>b</sup> Department of Earth Sciences, Durham University, Sciences Labs, Durham, DH1 3LE, United Kingdom

<sup>c</sup> School of Earth Sciences, University of Bristol, Bristol, BS8 1RJ, United Kingdom

<sup>d</sup> Department of Environment, Earth and Ecosystems, CEPESAR, The Open University, Walton Hall, Milton Keynes, MK7 6AA, United Kingdom

### ARTICLE INFO

#### Article history:

Received 16 November 2016

Accepted 17 November 2016

Available online 20 December 2016

Editor: B. Marty

The authors have corrected a few minor mistakes that appear in [Table 2](#), that is, four  $\epsilon^{54}\text{Cr}$  values have been updated (JP-1, Bremerförde, Kernouve and Saint-Severin). These minor changes do not affect the figures or the conclusions of the article.

The corrected [Table 2](#) appears below for the reader's convenience.

DOI of original article: <http://dx.doi.org/10.1016/j.epsl.2015.11.026>.

\* Corresponding author. Fax: +44 (0)1865 272072.

E-mail address: [pierre.bonnand@earth.ox.ac.uk](mailto:pierre.bonnand@earth.ox.ac.uk) (P. Bonnand).

<sup>1</sup> Present address: Department of Earth Sciences, University of Oxford, South Parks Road, Oxford, OX1 3AN, United Kingdom.

<http://dx.doi.org/10.1016/j.epsl.2016.11.029>

0012-821X/© 2016 The Author(s). Published by Elsevier B.V. This is an open access article under the CC BY license (<http://creativecommons.org/licenses/by/4.0/>).

**Table 2**

Mass independent and mass dependent Cr isotopic compositions for meteorites and terrestrial samples analysed in this study. The Cr concentrations are from Qin et al. (2010), Trinquier et al. (2008a) and Shukolyukov and Lugmair (2006).

Sample name	Group	Cr ( $\mu\text{g g}^{-1}$ )	$\epsilon^{53}\text{Cr}$	$2\sigma^{\text{a}}$	$\epsilon^{54}\text{Cr}$	$2\sigma^{\text{a}}$	$n^{\text{b}}$	$\delta^{53}\text{Cr}$ (‰)	$2\sigma^{\text{c}}$	$n^{\text{d}}$
NIST SRM979		n/a	0.00	0.10	0.00	0.19	15	0.000	0.009*	15
JP-1		2780	0.11	0.10	0.18	0.19	3	−0.102	0.012	5
<i>Carb. Chondrite</i>										
Allende	CV3	3600	0.14	0.10	1.04	0.19	1	−0.113	0.012	4
Orgueil	CI	2650	0.53	0.10	1.45	0.19	6	−0.145	0.012	6
Murchison	CM2	3100	0.36	0.10	0.98	0.19	1	−0.119	0.012	1
Ormans	CO3	3300	0.19	0.10	1.03	0.19	1	−0.068	0.012	1
Average CC								−0.111	0.065	
<i>Ord. Chondrite</i>										
Bremervörde	H/L3	3600	0.36	0.10	−0.11	0.19	1	−0.146	0.012	1
Kernouve	H6	3700	0.27	0.10	−0.21	0.19	1	−0.111	0.012	1
Barratta	L4	3740	0.29	0.10	−0.10	0.19	1	−0.118	0.012	1
Bruderheim	L6	3400	0.30	0.10	−0.26	0.19	1	−0.112	0.012	1
Parnallee	LL3	3200	0.38	0.10	−0.08	0.19	1	−0.117	0.012	1
Saint-Severin	LL6	3970	0.38	0.10	0.02	0.19	1	−0.104	0.012	1
Average OC								−0.118	0.029	
<i>Enst. Chondrite</i>										
Indarch	EH4	3400	0.37	0.10	0.15	0.19	1	−0.103	0.012	1
Khairpur	EL6	4060	0.27	0.10	0.14	0.19	1	−0.048	0.012	1
Average EC								−0.076	0.078	
<i>HED</i>										
Kapoeta	How.	3600	0.32	0.10	−0.71	0.19	1	−0.156	0.012	1
Pasamonte	Eucrite	1750	1.03	0.10	−0.45	0.19	1	−0.176	0.012	1
Juvinas	Eucrite	2100	0.93	0.10	−0.55	0.19	1	−0.184	0.012	1
Johnstown	Diog.	5700	0.32	0.10	−0.56	0.19	1	−0.074	0.012	1
<i>Mars</i>										
Nakhla	SNC	1770	0.27	0.10	−0.35	0.19	1	−0.156	0.012	1

Note: Carb.: carbonaceous, Ord.: ordinary, Enst.: enstatite, How.: howardite, Diog.: diogenite.

<sup>a</sup> 2 standard deviations of repeated measurements of the NIST 979 standard.

<sup>b</sup> Number of unspiked runs.

<sup>c</sup> 2 standard deviations of repeated measurements of Orgueil sample.

<sup>d</sup> Number of spiked runs.

\* 2 standard deviations of the NIST 979 (see text for details).